1985RMxAA..11..159M

ERRATA

In *Revista Mexicana Astron. Astrof.*, Vol. **10** (1985) the Index should be modified as follows:

Page 4, lines 37 and 38 : THE BEHAVIOUR OF THE ATMOSPHERE OF σ SCORPII,
A. Costa and A. Ringuelet, 293.

Page 5, lines 4 and 5 : ALGUMAS NOVAS PERSPECTIVAS EM MECANICA DO SISTEMA SOLAR*,
S. Ferraz-Mello, 339.

The paper: THE PROGRESSIVE OCCULTATION OF THE BINARY CENTRAL STAR OF NGC 2346 BY A DENSE DUST CLOUD*, by R.H. Méndez, B.F. Marino, J.J. Clariá, and W. van Driel, (*Revista Mexicana Astron. Astrof.*, Vol. **10**, 187-197, 1985) should be ammended to read as follows:

Page 189, TABLE 2 : star c has $U-B=0.15$

Page 193, line 6 from the bottom: "... to the A-type star, its brightness should be maximum at the time when the brightness of the A-type star is minimum, that is to say near orbital phase 0.8 ..."

Page 194, line 14: "... component of the orbit's motion must be substantially larger than the vertical component, for ..."

Page 194, line 23: "... more than 10^7 km. We shall call this distance ΔX (see Fig. 4, which is explained below). ..."

Page 194, line 24: "... We can obtain the corresponding column densities N_d (dust particles/cm^2) from ..."

Page 195, Caption to Figure 4: "... to minimum visual absorption produced by the cloud (see text and and Table 5 ..."

En el artículo: CINEMATICA GALACTICA LOCAL: UN MODELO ISOTERMICO, por Jorge Núñez (*Revista Mexicana Astron. Astrof.*, Vol. **8**, 91-108 (1983), la Tabla 2, Muestra 4 debe de modificarse de la siguiente manera:

TABLA 2

RESULTADOS DE LA RESOLUCION POR MINIMOS ....

Muestra 4

X	0.02	0.02	0.04	0.02	-0.01	0.01	0.01	0.01
Y	-0.41	0.01	-0.35	0.02	-0.34	0.01	-0.37	0.01
Z	0.26	0.02	0.16	0.01	0.15	0.01	0.20	0.01
$S_{11}-S_{33}$	-0.31	0.27	-0.34	0.35	-0.20	0.15	-0.16	0.13
S_{12}	0.24	0.08	0.20	0.11	0.21	0.05	0.20	0.04
S_{13}	-0.02	0.12	-0.40	0.16	0.02	0.08	0.01	0.07
$S_{22}-S_{33}$	-0.36	0.27	0.55	0.28	0.39	0.15	0.03	0.13
S_{23}	-0.27	0.13	-0.33	0.18	0.03	0.08	-0.10	0.07
S_{33}	0.65	0.23	----	----	----	----	0.46	0.12
ω_1	----	----	0.20	0.07	0.27	0.05	0.27	0.05
ω_2	----	----	-0.52	0.11	-0.41	0.07	-0.53	0.07
ω_3	----	----	----	----	-0.37	0.06	-0.42	0.06

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